
Mathematical thermodynamics of viscous fluids

Eduard Feireisl¹

Institute of Mathematics, Czech Academy of Sciences, Praha
feireisl@math.cas.cz

Abstract. This course is a short introduction to the mathematical theory of the motion of viscous fluids. We introduce the concept of weak solution to the Navier-Stokes-Fourier system and discuss its basic properties. In particular, we construct the weak solutions as a suitable limit of a mixed numerical scheme based on a combination of the finite volume and finite elements method. The question of stability and robustness of various classes of solutions is addressed with the help of the relative (modulated) energy functional. Related results concerning weak-strong uniqueness and conditional regularity of weak solutions are presented. Finally, we discuss the asymptotic limit when viscosity of the fluid tends to zero. Several examples of ill-posedness for the limit Euler system are given and an admissibility criterion based on the viscous approximation is proposed.

Motto: *Die Energie der Welt ist constant; Die Entropie der Welt strebt einem Maximum zu.*

Rudolph Clausius (1822–1888)

1 Fluids in continuum mechanics

We start by introducing a mathematical structure of the theory of *fluids* in the framework of continuum mechanics.

1.1 Fluids in equilibrium

A fluid in equilibrium is characterized by two basic state variables: the mass density ϱ and the (specific) internal energy e . Following Callen [5] we introduce the specific entropy s , which is a function of ϱ , e enjoying the following properties:

1. The entropy s is an increasing function of the internal energy,

$$\frac{\partial s}{\partial e} = \frac{1}{\vartheta}, \quad \vartheta > 0,$$

where ϑ is the absolute temperature.

2. For a thermally and mechanically insulated fluid occupying a physical domain Ω , maximization of the total entropy

$$S = \int_{\Omega} \varrho s \, dx$$

yields the equilibrium state of the system.

3. Third law of thermodynamics:

$$s \rightarrow 0 \text{ as } \vartheta \rightarrow 0.$$

In what follows, it will be convenient to consider ϱ and ϑ as the basic variables characterizing the state of a fluid, the other thermodynamic quantities e , s are interrelated by means of *Gibbs' equation*:

$$\vartheta Ds = De + pD\left(\frac{1}{\varrho}\right), \quad (1)$$

where p is the pressure.

1.2 Fluids in motion

Suppose a fluid occupies a part of the physical space R^3 represented by a domain Ω . We adopt the Eulerian description of motion taking the coordinate system attached to Ω rather than to the fluid itself. The motion is characterized by the macroscopic velocity $\mathbf{u} = \mathbf{u}(t, x)$ - a function of the time $t \in (0, T)$ and $x \in \Omega$. The streamlines \mathbf{X} are determined by the system of ordinary differential equations

$$\frac{d\mathbf{X}}{dt} = \mathbf{u}(t, \mathbf{X}), \quad \mathbf{X}(0) = x \in \Omega. \quad (2)$$

1.3 Field equations

A suitable mathematical description of fluids in continuum mechanics is given by a system of field equations expressing the basic physical principles.

Mass conservation

Mass conservation in fluid dynamics is formulated through *equation of continuity*

$$\partial_t \varrho + \operatorname{div}_x(\varrho \mathbf{u}) = 0, \quad (3)$$

or, if the velocity field is smooth,

$$\partial_t \varrho + \mathbf{u} \cdot \nabla_x \varrho = -\varrho \operatorname{div}_x \mathbf{u}, \quad (4)$$

where the left-hand side

$$\partial_t \varrho + \mathbf{u} \cdot \nabla_x \varrho = \frac{d}{dt} \varrho(t, \mathbf{X}(t))$$

describes the mass transport along the stream-lines while the “source” term represents its changes due to compressibility.

Momentum balance

Momentum balance is enforced through *Newton’s second law*. Introducing the Cauchy stress tensor \mathbb{T} we get

$$\partial_t(\varrho \mathbf{u}) + \operatorname{div}_x(\varrho \mathbf{u} \otimes \mathbf{u}) = \operatorname{div}_x \mathbb{T} + \varrho \mathbf{f}, \quad (5)$$

where \mathbf{f} is the mass density of the external volume forces that may be acting on the fluid.

Fluids as materials in continuum mechanics are characterized by *Stokes law*:

$$\mathbb{T} = \mathbb{S} - p\mathbb{I}, \quad (6)$$

where \mathbb{S} is the viscous stress tensor, the basic properties of which will be discussed below.

Energy and entropy

The kinetic energy balance is obtained by taking the scalar product of the velocity with the momentum equation (5):

$$\begin{aligned} \partial_t \left(\frac{1}{2} \varrho |\mathbf{u}|^2 \right) + \operatorname{div}_x \left[\left(\frac{1}{2} \varrho |\mathbf{u}|^2 + p \right) \mathbf{u} \right] + \operatorname{div}_x (\mathbb{S} \cdot \mathbf{u}) \\ = p \operatorname{div}_x \mathbf{u} - \mathbb{S} : \nabla_x \mathbf{u} + \varrho \mathbf{f} \cdot \mathbf{u}. \end{aligned} \quad (7)$$

Accordingly, the internal energy equation takes the form

$$\partial_t \varrho e + \operatorname{div}_x(\varrho e \mathbf{u}) + \operatorname{div}_x \mathbf{q} = \mathbb{S} : \nabla_x \mathbf{u} - p \operatorname{div}_x \mathbf{u}, \quad (8)$$

where \mathbf{q} denotes the heat flux, whereas the total energy balance expressing the *First law of thermodynamics* reads

$$\partial_t \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e \right) + \operatorname{div}_x \left[\left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e + p \right) \mathbf{u} \right] + \operatorname{div}_x (\mathbb{S} \cdot \mathbf{u} + \mathbf{q}) = \varrho \mathbf{f} \cdot \mathbf{u}. \quad (9)$$

Here, we have deliberately omitted the effect of external heat (energy) sources, and, hereafter, we shall also ignore the external force \mathbf{f} unless specified otherwise.

There several alternative ways how to express the total energy balance, all of the equivalent to (9) within the class of *smooth* solutions. Introducing the thermal pressure p_ϑ and the specific heat at constant volume c_v ,

$$p_\vartheta = \frac{\partial p(\varrho, \vartheta)}{\partial \vartheta}, \quad c_v = \frac{\partial e(\varrho, \vartheta)}{\partial \vartheta},$$

we may use Gibbs' relation (1) to deduce an alternative formulation of (8) frequently used in the literature:

$$\varrho c_v (\partial_t \vartheta + \mathbf{u} \cdot \nabla_x \vartheta) + \operatorname{div}_x \mathbf{q} = \mathbb{S} : \nabla_x \mathbf{u} - \vartheta p_\vartheta \operatorname{div}_x \mathbf{u}. \quad (10)$$

Another consequence of (1) is the entropy balance

$$\partial_t (\varrho s) + \operatorname{div}_x (\varrho s \mathbf{u}) + \operatorname{div}_x \left(\frac{\mathbf{q}}{\vartheta} \right) = \sigma, \quad \sigma = \frac{1}{\vartheta} \left(\mathbb{S} : \nabla_x \mathbf{u} - \frac{\mathbf{q} \cdot \nabla_x \vartheta}{\vartheta} \right) \quad (11)$$

that may be seen as a mathematical formulation of the Second law of thermodynamics. Accordingly, the entropy production rate σ must be non-negative for any physically admissible process yielding the restriction

$$\mathbb{S} : \nabla_x \mathbf{u} \geq 0, \quad -\mathbf{q} \cdot \nabla_x \vartheta \geq 0. \quad (12)$$

1.4 Boundary behavior

Any real physical domain Ω is bounded although some problems may be conveniently posed on unbounded domains. In both cases, the boundary behavior of the fluid is relevant for determining the motion inside Ω . We focus on very simple boundary conditions yielding energetically insulated fluid systems. Specifically, we impose the impermeability of the boundary

$$\mathbf{u} \cdot \mathbf{n}|_{\partial\Omega} = 0, \quad \mathbf{n} \text{ -the outer normal vector to } \partial\Omega, \quad (13)$$

supplemented, in the case of viscous fluids, with either the no-slip

$$\mathbf{u} \times \mathbf{n}|_{\partial\Omega} = 0, \quad (14)$$

or the complete slip

$$[\mathbb{S} \cdot \mathbf{n}] \times \mathbf{n}|_{\partial\Omega} = 0. \quad (15)$$

In addition, the no-flux boundary conditions will be imposed on the heat flux

$$\mathbf{q} \cdot \mathbf{n}|_{\partial\Omega} = 0. \quad (16)$$

Accordingly, in the absence of external forces, the total mass as well as the total energy are conserved quantities:

$$\frac{d}{dt} \int_{\Omega} \varrho \, dx = 0, \quad \frac{d}{dt} \int_{\Omega} \left[\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e(\varrho, \vartheta) \right] dx = 0. \quad (17)$$

2 Mathematics of viscous compressible fluids

In this section, we introduce the concept of *weak solution* to the system of fields equations. To begin, the initial state of the system will be prescribed:

$$\varrho(0, \cdot) = \varrho_0, \vartheta(0, \cdot) = \vartheta_0, \mathbf{u}(0, \cdot) = \mathbf{u}_0, \varrho_0 > 0, \vartheta_0 > 0. \quad (18)$$

2.1 Equation of continuity

We say that ϱ, \mathbf{u} satisfy equation (5) in a weak sense if

$$\int_{\tau_1}^{\tau_2} \int_{\Omega} [\varrho \partial_t \varphi + \varrho \mathbf{u} \cdot \nabla_x \varphi] \, dx dt = \left[\int_{\Omega} \varrho \varphi \, dx \right]_{t=\tau_1}^{t=\tau_2} \quad (19)$$

for any $0 \leq \tau_1 \leq \tau_2 \leq T$ and any $\varphi \in C^\infty([0, T] \times \overline{\Omega})$.

We also introduce the *renormalized solutions* satisfying

$$\begin{aligned} \int_{\tau_1}^{\tau_2} \int_{\Omega} [b(\varrho) \partial_t \varphi + b(\varrho) \mathbf{u} \cdot \nabla_x \varphi + (b(\varrho) - b'(\varrho) \varrho) \operatorname{div}_x \mathbf{u}] \, dx dt \\ = \left[\int_{\Omega} b(\varrho) \varphi \, dx \right]_{t=\tau_1}^{t=\tau_2} \end{aligned} \quad (20)$$

for any smooth function b satisfying suitable growth conditions. Renormalized solutions to transport equations were introduced in the seminal paper by DiPerna and Lions [15].

Note that (19) can be seen as a special case of (20) with $b(\varrho) = \varrho$. Both (19) and (20) require certain *continuity* of ϱ as a function of the time.

2.2 Momentum equation

A weak formulation of the momentum balance (5), (6) (with $\mathbf{f} = 0$) reads

$$\begin{aligned} \int_{\tau_1}^{\tau_2} \int_{\Omega} [\varrho \mathbf{u} \cdot \varphi + \varrho \mathbf{u} \otimes \mathbf{u} : \nabla_x \varphi + p \operatorname{div}_x \varphi - \mathbb{S} : \nabla_x \varphi] \, dx dt \\ \left[\int_{\Omega} \varrho \mathbf{u} \cdot \varphi \, dx \right]_{t=\tau_1}^{t=\tau_2} \end{aligned} \quad (21)$$

for any $0 \leq \tau_1 \leq \tau_2 \leq T$ and any test function $\varphi \in C^\infty([0, T] \times \overline{\Omega}; \mathbb{R}^3)$. In addition, we require $\varphi \in C_c^\infty([0, T] \times \Omega)$ in the case of the no-slip boundary conditions (13), (14), and $\varphi \cdot \mathbf{n}|_{\partial\Omega} = 0$ for the complete slip (13), (15).

2.3 Energy-entropy

In order to close the system of field equations a weak formulation of the energy and/or entropy is needed. Accordingly, we have to choose *one* among the equations (9–11) as a suitable “representative” keeping in mind they are not equivalent in the weak setting.

Entropy based weak formulation

Of course, the most natural candidate would be the total energy balance (9) expressing the First law. In view of the technical difficulties discussed later in this text, however, we opted for the Second law encoded in (11), where, in addition, we allow the entropy production rate σ to be a non-negative measure satisfying

$$\sigma \geq \frac{1}{\vartheta} \left(\mathbb{S} : \nabla_x \mathbf{u} - \frac{\mathbf{q} \cdot \nabla_x \vartheta}{\vartheta} \right). \quad (22)$$

Accordingly, a weak formulation of (11), (22) reads

$$\begin{aligned} & \int_{\tau_1}^{\tau_2} \int_{\Omega} \left[\varrho s \partial_t \varphi + \varrho s \mathbf{u} \cdot \nabla_x \varphi + \frac{\mathbf{q}}{\vartheta} \cdot \nabla_x \varphi \right] dx dt \\ & + \int_{\tau_1}^{\tau_2} \int_{\Omega} \frac{1}{\vartheta} \left[\mathbb{S} : \nabla_x \mathbf{u} - \frac{\mathbf{q} \cdot \nabla_x \vartheta}{\vartheta} \right] \varphi dx dt \leq \left[\int_{\Omega} \varrho s \varphi dx \right]_{t=\tau_1}^{t=\tau_2} \end{aligned} \quad (23)$$

for a.a. $0 \leq \tau_1 \leq \tau_2 \leq T$ including $\tau_1 = 0$, and any $\varphi \in C^\infty([0, T] \times \overline{\Omega})$, $\varphi \geq 0$.

Replacing *equation* by *inequality* may result in a lost of information that must be compensated by adding an extra constraint. Here, we supplement (23) with the total energy balance

$$\left[\int_{\Omega} \left[\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e(\varrho, \vartheta) \right] dx \right]_{t=\tau_1}^{t=\tau_2} = 0 \quad (24)$$

for a.a. $0 \leq \tau_1 \leq \tau_2 \leq T$, including $\tau_1 = 0$.

Although the entropy formulation may seem rather awkward and unnecessarily complicated, it turns out to be quite convenient to deal with in the framework of weak solutions. In particular, it gives rise to the relative energy inequality with the associated concept of dissipative solution discussed below.

Thermal energy weak formulation

If c_v is constant, we may replace (23) by the thermal energy balance (10) written again as an inequality

$$\begin{aligned} & \int_{\tau_1}^{\tau_2} \int_{\Omega} \left[c_v \varrho \vartheta \partial_t \varphi + c_v \varrho \vartheta \mathbf{u} \cdot \nabla_x \varphi + \mathbf{q} \cdot \nabla_x \varphi \right] dx dt \\ & + \int_{\tau_1}^{\tau_2} \int_{\Omega} \left[\mathbb{S} : \nabla_x \mathbf{u} - \varrho p_\vartheta \operatorname{div}_x \mathbf{u} \right] \varphi dx dt \leq \left[\int_{\Omega} c_v \varrho \vartheta \varphi dx \right]_{t=\tau_1}^{t=\tau_2} \end{aligned} \quad (25)$$

for a.a. $0 \leq \tau_1 \leq \tau_2 \leq T$ including $\tau_1 = 0$, and any $\varphi \in C^\infty([0, T] \times \overline{\Omega})$, $\varphi \geq 0$. Similarly to the preceding section, the inequality (25) is supplemented by the total energy balance

$$\left[\int_{\Omega} \left[\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e(\varrho, \vartheta) \right] dx \right]_{t=\tau_1}^{t=\tau_2} = 0 \quad (26)$$

for a.a. $0 \leq \tau_1 \leq \tau_2 \leq T$, including $\tau_1 = 0$.

The weak formulation based on the thermal energy balance is simpler than (23), (24) and can be used in the analysis of the associated *numerical schemes*.

2.4 Constitutive relations, Navier-Stokes-Fourier system

In order to obtain, at least formally, a mathematically well-posed problem, the constitutive relations for the viscous stress tensor \mathbb{S} as well as the heat flux \mathbf{q} must be specified. They must obey the general principle (12) in order to comply with the Second law of thermodynamics.

We consider the simplest possible situation when \mathbb{S} is a linear function of the velocity gradient $\nabla_x \mathbf{u}$, while \mathbf{q} depends linearly on $\nabla_x \vartheta$. More specifically, we impose Newton's rheological law

$$\mathbb{S}(\vartheta, \nabla_x \mathbf{u}) = \mu(\vartheta) \left[\nabla_x \mathbf{u} + \nabla_x^t \mathbf{u} - \frac{2}{3} \operatorname{div}_x \mathbf{u} \mathbb{I} \right] + \eta(\vartheta) \operatorname{div}_x \mathbf{u} \mathbb{I}, \quad (27)$$

where $\mu(\vartheta)$, $\eta(\vartheta)$ are non-negative scalar functions representing the shear and bulk viscosity coefficient, respectively.

Similarly, the heat flux will obey Fourier's law

$$\mathbf{q}(\vartheta, \nabla_x \vartheta) = -\kappa(\vartheta) \nabla_x \vartheta, \quad (28)$$

with the heat conductivity coefficient κ .

The system of field equations supplemented with the constitutive relations (27), (28) will be referred to as *Navier-Stokes-Fourier system*.

3 Well-posedness, approximation scheme

The question of *well-posedness* of a system of equations, including the problem of existence, uniqueness, and stability with respect to the data, is crucial in the mathematical theory. As is well known, well posedness for the equations and systems arising in fluid mechanics features sofar unsurmountable mathematical difficulties due to the occurrence of possible singularities, in particular in the velocity field (cf. Fefferman [19]). From this point of view, the concept of weak solution offers a suitable framework to attack the problem. Moreover, the weak solutions seem indispensable in the theory of *inviscid fluids* where the singularities *are known* to occur in finite time no matter how smooth the initial data are.

Although the question of mere *existence* of solutions to problems like the Navier-Stokes-Fourier system seems extremely difficult, a more ambitious task

is to design a suitable *approximation scheme* usable in effective numerical implementations. In accordance with the philosophy proposed in the nowadays classical monograph by J.-L.Lions [33], solutions should be obtained as limits of a finite number of algebraic equations solvable by means of a suitable numerical method.

3.1 An approximation scheme for the Navier-Stokes-Fourier system

We propose a discrete approximation scheme for solving the Navier-Stokes-Fourie system. To this end, we employ the thermal energy formulation based on the relations (19), (21), (25), and (26).

Hypotheses

Several mostly technical restrictions must be imposed on the constitutive equations to make the problem tractable by means of the available analytical tools. Specifically, we suppose that:

- The internal energy $e(\varrho, \vartheta)$ can be written in the form $e(\varrho, \vartheta) = c_v \vartheta + P(\varrho)$, where the specific heat at constant volume c_v is a positive constant. Accordingly, we suppose that the pressure takes the form

$$p(\varrho, \vartheta) = a\varrho^\gamma + b\varrho + \varrho\vartheta, \quad a, b > 0, \quad \gamma > 3, \quad (29)$$

therefore the (specific) internal energy reads

$$e(\varrho, \vartheta) = c_v \vartheta + P(\varrho), \quad c_v > 0, \quad P(\varrho) = \frac{a}{\gamma - 1} \varrho^\gamma + b\varrho \log(\varrho). \quad (30)$$

- The viscosity coefficients in (27) are constant, in particular, we may write

$$\operatorname{div}_x \mathbb{S}(\nabla_x \mathbf{u}) = \mu \Delta \mathbf{u} + \lambda \nabla_x \operatorname{div}_x \mathbf{u}, \quad \lambda = \frac{1}{3} \mu + \eta > 0. \quad (31)$$

- The heat flux \mathbf{q} obeys Fourier's law (28) where the heat conductivity coefficient κ is a continuously differentiable function of the temperature satisfying

$$\kappa = \kappa(\vartheta), \quad \underline{\kappa}(1 + \vartheta^2) \leq \kappa(\vartheta) \leq \bar{\kappa}(1 + \vartheta^2), \quad \underline{\kappa} > 0. \quad (32)$$

3.2 Time discretization

We propose to approach the Navier-Stokes-Fourier by Rothe's method or the method of *time discretization*. We fix the time step $\Delta t > 0$, and, supposing the approximate solutions $[\varrho^j, \vartheta^j, \mathbf{u}^j]$ at the times $j\Delta t$, $j = 0, \dots, k-1$ already determined by previous steps, we define $[\varrho^k, \vartheta^k, \mathbf{u}^k]$ as a solution of the system of "stationary" problems

$$\begin{aligned}
 D_t \varrho^k &\equiv \frac{\varrho^k - \varrho^{k-1}}{\Delta t} = \mathcal{C}(\varrho^k, \vartheta^k, \mathbf{u}^k), \\
 D_t(\varrho^k \mathbf{u}^k) &\equiv \frac{\varrho^k \mathbf{u}^k - \varrho^{k-1} \mathbf{u}^{k-1}}{\Delta t} = \mathcal{M}(\varrho^k, \vartheta^k, \mathbf{u}^k), \\
 c_v D_t(\varrho^k \vartheta^k) &\equiv c_v \frac{\varrho^k \vartheta^k - \varrho^{k-1} \vartheta^{k-1}}{\Delta t} = \mathcal{T}(\varrho^k, \vartheta^k, \mathbf{u}^k),
 \end{aligned} \tag{33}$$

for certain operators \mathcal{C} , \mathcal{M} , and \mathcal{T} . Such a scheme is called *implicit* as we have to solve a system of (non-linear) equations to determine $[\varrho^k, \vartheta^k, \mathbf{u}^k]$ at each time step.

3.3 Space discretization

Our goal is to approximate each equation in (33) by a finite system of algebraic equations. This is usually done by applying a suitable projection onto a finite dimensional space. Accordingly, we replace $\varrho^k \approx \varrho_h^k$, $\vartheta \approx \vartheta_h^k$, $\mathbf{u} \approx \mathbf{u}_h^k$ by finite vectors, where the parameter $h > 0$ characterizes the degree of space discretization.

Similarly to the weak formulation discussed in Section 2, we multiply the corresponding equations by a test function belonging to a suitable *finite-dimensional* space and perform the by-parts integration. In such a way, we may for instance replace the differential operator

$$-\operatorname{div}_x \mathbb{S}(\nabla_x \mathbf{u}^k) \approx \mu \int_{\Omega} \nabla_x \mathbf{u}^k : \nabla_x \phi \, dx + \lambda \int_{\Omega} \operatorname{div}_x \mathbf{u}_h^k \operatorname{div}_x \phi \, dx.$$

In numerical analysis, such a step is usual performed via a *finite element* method (FEM). The physical domain Ω is approximated by a numerical domain Ω_h , the latter being divided into small elementary pieces by *triangulation*. The test functions restricted to these elementary pieces (elements) are usually polynomials of finite degree enjoying certain continuity on the faces common to two neighboring elements. In order to specify our approximation scheme we start by a short excursion in numerical analysis.

Mesh, triangulation

The physical space Ω is approximated by a *polyhedral bounded domain* $\Omega_h \subset \mathbb{R}^3$ that admits a *tetrahedral* mesh E_h ; the individual elements in the mesh will be denoted by $E \in E_h$. Faces in the mesh are denoted as Γ , whereas Γ_h is the set of all faces. Moreover, the set of faces $\Gamma \subset \partial\Omega_h$ is denoted $\Gamma_{h,\text{ext}}$, while $\Gamma_{h,\text{int}} = \Gamma_h \setminus \Gamma_{h,\text{ext}}$. The size (diameter h_E of its elements E in the mesh) is proportional to a positive parameter h .

In addition, the mesh enjoys certain additional properties (cf. Eymard et al. [17, Chapter 3]):

- The intersection $E \cap F$ of two elements $E, F \in E_h$, $E \neq F$ is either empty or their common face, edge, or vertex.

- For any $E \in E_h$, $\text{diam}[E] \approx h$, $r[E] \approx h$, where r denotes the radius of the largest sphere contained in E .
- There is a family of control points $x_E \in E$, $E \in E_h$ such that if E and F are two neighboring elements sharing a common face Γ , then the segment $[x_E, x_F]$ is perpendicular to Γ . We denote

$$d_\Gamma = |x_E - x_F| > 0.$$

Each face $\Gamma \in \Gamma_h$ is associated with a normal vector \mathbf{n} . We shall write Γ_E whenever a face $\Gamma_E \subset \partial E$ is considered as a part of the boundary of the element E . In such a case, the normal vector to Γ_E is always the *outer* normal vector with respect to E . Moreover, for any function g continuous on each element E , we set

$$g^{\text{out}}|_\Gamma = \lim_{\delta \rightarrow 0^+} g(\cdot + \delta \mathbf{n}), \quad g^{\text{in}}|_\Gamma = \lim_{\delta \rightarrow 0^+} g(\cdot - \delta \mathbf{n}),$$

$$[[g]]_\Gamma = g^{\text{out}} - g^{\text{in}}, \quad \{g\}_\Gamma = \frac{1}{2} (g^{\text{out}} + g^{\text{in}}).$$

For $\Gamma_E \subset \partial E$ we simply write g for g^{in} . We also omit the subscript Γ if no confusion arises.

Finally, we distinguish two families of faces,

$$\Gamma_{h,\text{ext}} = \left\{ \Gamma \in \Gamma_h \mid \Gamma \subset \partial \Omega_h \right\}, \quad \Gamma_{h,\text{int}} = \Gamma_h \setminus \Gamma_{h,\text{ext}}.$$

FEM structure

The velocity field \mathbf{u}^k will be approximated by the so-called Crouzeix-Raviart finite elements (see for instance Brezzi and Fortin [4]) belonging to the space

$$V_h(\Omega_h) = \left\{ v \in L^2(\Omega_h) \mid v|_E = \text{affine function}, \quad E \in E_h, \right. \\ \left. \int_\Gamma [[v]] \, dS_x = 0 \text{ for any } \Gamma \in \Gamma_{h,\text{int}} \right\}.$$

For the sake of definiteness, we focus on the no-slip boundary conditions (13), (14). Accordingly, we introduce the space

$$V_{h,0}(\Omega_h) = \left\{ v \in V_h \mid \int_\Gamma v \, dS_x = 0 \text{ for any } \Gamma \in \Gamma_{h,\text{ext}} \right\}.$$

The finite element method is suitable for approximating the viscous stress, however, the Navier-Stokes-Fourier system contains also hyperbolic like equations as (3). The convective terms, appearing in any of the field equations are better approximated by means of the finite volume method (FVM) introduced in the next section.

FVM structure

Roughly speaking, the finite volume method replaces integration over the elements by integration over faces. To this end, we introduce the space of piece-wise constant functions

$$Q_h(\Omega_h) = \left\{ v \in L^2(\Omega_h) \mid v|_E = a_E \in R \text{ for any } E \in E_h \right\},$$

along with the associated projection operator

$$\Pi_h^Q : L^1(\Omega_h) \rightarrow Q_h(\Omega_h), \quad \Pi_h^Q[v] \equiv \hat{v}, \quad \Pi_h^Q[v]|_E = \frac{1}{|E|} \int_E v \, dx \text{ for any } E \in E_h.$$

The convective terms are discretized by means of the so-called upwind defined as follows:

$$\text{Up}[r, \mathbf{u}] = r^{\text{in}}[\tilde{\mathbf{u}} \cdot \mathbf{n}]^+ + r^{\text{out}}[\tilde{\mathbf{u}} \cdot \mathbf{n}]^-,$$

where we have denoted

$$[c]^+ = \max\{c, 0\}, \quad [c]^- = \min\{c, 0\}, \quad \tilde{v} = \frac{1}{|\Gamma|} \int_\Gamma v \, dS_x.$$

Such a definition makes sense as soon as $r \in Q_h(\Omega_h)$, $\mathbf{u} \in V_h(\Omega_h; R^3)$ and $\Gamma \in \Gamma_{h,\text{int}}$. We approximate

$$\int_{\Omega_h} \varrho Z \mathbf{u} \cdot \nabla_x \phi \, dx \approx \sum_{\Gamma \in \Gamma_{h,\text{int}}} \int_\Gamma \text{Up}[\varrho Z, \mathbf{u}] [[\phi]] \, dS_x.$$

3.4 Approximation scheme

We introduce the notation ∇_h , div_h to denote the restriction of the differential operators ∇_x , div_x to any element $E \in E_h$, specifically

$$\nabla_h v|_E = \nabla_x v|_E, \quad \text{div}_h \mathbf{v}|_E = \text{div}_x \mathbf{v}|_E$$

for functions v , \mathbf{v} differentiable on any $E \in E_h$.

The zero-th order terms $[\varrho^0, \vartheta^0, \mathbf{u}^0]$ being determined by the initial data, we define $[\varrho_h^k, \vartheta_h^k, \mathbf{u}_h^k]$ successively as a solution to the approximation scheme - numerical method:

$$\begin{aligned} \int_{\Omega_h} D_t \varrho_h^k \phi \, dx - \sum_{\Gamma \in \Gamma_{h,\text{int}}} \int_\Gamma \text{Up}[\varrho_h^k, \mathbf{u}_h^k] [[\phi]] \, dS_x & \quad (34) \\ + h^\alpha \sum_{\Gamma \in \Gamma_{h,\text{int}}} \int_\Gamma [[[\varrho_h^k]]] [[\phi]] \, dS_x & = 0 \end{aligned}$$

for all $\phi \in Q_h(\Omega_h)$, with a parameter $0 < \alpha < 1$;

$$\begin{aligned} & \int_{\Omega_h} D_t(\varrho_h^k \widehat{\mathbf{u}}_h^k) \cdot \phi \, dx - \sum_{\Gamma \in \Gamma_{h,\text{int}}} \int_{\Gamma} \text{UP}[\varrho_h^k \widehat{\mathbf{u}}_h^k, \mathbf{u}_h^k] \cdot \left[\left[\widehat{\phi} \right] \right] \, dS_x \quad (35) \\ & + \int_{\Omega_h} [\mu \nabla_h \mathbf{u}_h^k : \nabla_h \phi + \lambda \text{div}_h \mathbf{u}_h^k \text{div}_h \phi] \, dx - \int_{\Omega_h} p(\varrho_h^k, \vartheta_h^k) \text{div}_h \phi \, dx \\ & \quad + h^\alpha \sum_{\Gamma \in \Gamma_{h,\text{int}}} \int_{\Gamma} \left[\left[\varrho_h^k \right] \right] \left\{ \widehat{u}_h^k \right\} \cdot \left[\left[\widehat{\phi} \right] \right] \, dS_x = 0 \end{aligned}$$

for any $\phi \in V_{h,0}(\Omega_h; \mathbb{R}^3)$;

$$\begin{aligned} & c_v \int_{\Omega_h} D_t(\varrho_h^k \vartheta_h^k) \phi \, dx - c_v \sum_{\Gamma \in \Gamma_{h,\text{int}}} \int_{\Gamma} \text{UP}[\varrho_h^k \vartheta_h^k, \mathbf{u}_h^k] \left[\left[\phi \right] \right] \, dS_x \quad (36) \\ & \quad + \sum_{\Gamma \in \Gamma_{h,\text{int}}} \int_{\Gamma} \frac{1}{d\Gamma} \left[\left[K(\vartheta_h^k) \right] \right] \left[\left[\phi \right] \right] \, dS_x \\ & = \int_{\Omega_h} [\mu |\nabla_h \mathbf{u}_h^k|^2 + \lambda |\text{div}_h \mathbf{u}_h^k|^2] \phi \, dx - \int_{\Omega_h} \varrho_h^k \vartheta_h^k \text{div}_h \mathbf{u}_h^k \phi \, dx \end{aligned}$$

for any $\phi \in Q_h(\Omega_h)$, where

$$K(\vartheta) = \int_0^\vartheta \kappa(z) \, dz.$$

Here, the terms proportional to h^α represent numerical counterparts of the artificial viscosity regularization used in [21, Chapter 7] and were introduced by Eymard et al. [18] to prove convergence of the momentum method.

3.5 Existence of weak solutions via the numerical scheme

The physical domain Ω will be approximated by the polyhedral domains Ω_h so that

$$\Omega \subset \overline{\Omega} \subset \Omega_h \subset \left\{ x \in \mathbb{R}^3 \mid \text{dist}[x, \overline{\Omega}] < h \right\}. \quad (37)$$

The discrete solutions $[\varrho_h^k, \vartheta_h^k, \mathbf{u}_h^k]$ can be extended to be defined at any time t setting

$$\begin{aligned} & \varrho_h(t, \cdot) = \varrho_h^0, \quad \vartheta_h(t, \cdot) = \vartheta_h^0, \quad \mathbf{u}_h(t, \cdot) = \mathbf{u}_h^0 \quad \text{for } t \leq 0, \\ & \varrho_h(t, \cdot) = \varrho_h^k, \quad \vartheta_h(t, \cdot) = \vartheta_h^k, \quad \mathbf{u}_h(t, \cdot) = \mathbf{u}_h^k \quad \text{for } t \in [k\Delta t, (k+1)\Delta t), \quad k = 1, 2, \dots, \end{aligned}$$

and, accordingly, the discrete time derivative of a quantity v_h is

$$D_t v_h(t, \cdot) = \frac{v_h(t) - v_h(t - \Delta t)}{\Delta t}, \quad t > 0.$$

We claim the following result:

Theorem 1. *Let $\Omega \subset R^3$ be a bounded domain of class C^1 approximated by a family of polyhedral domains $\{\Omega_h\}_{h>0}$ in the sense (37), where each Ω_h admits a tetrahedral mesh satisfying the hypotheses specified in Section 3.3. Suppose that $\mu > 0$, $\lambda > 0$, and that the pressure $p = p(\varrho, \vartheta)$ and the heat conductivity coefficient $\kappa = \kappa(\vartheta)$ comply with (29–32). Suppose that*

$$\Delta t \approx h$$

and

$$\varrho_h^0 > 0, \vartheta_h^0 > 0 \text{ for all } h > 0.$$

Then

- the numerical scheme (34), (35), (36) admits a solution

$$\varrho_h^k > 0, \vartheta_h^k > 0, \mathbf{u}_h^k > 0 \text{ for any finite } k = 1, 2, \dots;$$

-

$\varrho_h \rightarrow \varrho$ weakly-(*) in $L^\infty(0, T; L^\gamma(\Omega))$ and strongly in $L^1((0, T) \times \Omega)$,

$$\vartheta_h \rightarrow \vartheta \text{ weakly in } L^2(0, T; L^6(\Omega)),$$

$$\mathbf{u}_h \rightarrow \mathbf{u} \text{ weakly in } L^2(0, T; L^6(\Omega; R^3)),$$

$$\nabla_h \mathbf{u}_h \rightarrow \nabla_x \mathbf{u} \text{ weakly in } L^2((0, T) \times \Omega; R^{3 \times 3}),$$

at least for a suitable subsequence, where $[\varrho, \vartheta, \mathbf{u}]$ is a weak solution of the thermal energy formulation of the Navier-Stokes-Fourier system (19), (21), (25), (26) in $(0, T) \times \Omega$.

Remark 1. As a matter of fact, the thermal energy balance (25) is satisfied in the following very weak sense:

$$\begin{aligned} & \int_{\tau_1}^{\tau_2} \int_{\Omega} \left[c_v \varrho \vartheta \partial_t \varphi + c_v \varrho \vartheta \mathbf{u} \cdot \nabla_x \varphi - \overline{K(\vartheta)} \Delta \varphi \right] dx dt \\ & + \int_{\tau_1}^{\tau_2} \int_{\Omega} [\mathbb{S} : \nabla_x \mathbf{u} - \varrho p_{\vartheta} \operatorname{div}_x \mathbf{u}] \varphi dx dt \leq \left[\int_{\Omega} c_v \varrho \vartheta \varphi dx \right]_{t=\tau_1}^{t=\tau_2} \end{aligned}$$

for a.a. $0 \leq \tau_1 \leq \tau_2 \leq T$ including $\tau_1 = 0$, and any $\varphi \in C^\infty([0, T] \times \overline{\Omega})$, $\varphi \geq 0$, where

$$\overline{\varrho K(\vartheta)} = \varrho K(\vartheta).$$

The existence part in Theorem 1 was established in [24, Section 8.1], convergence of the numerical solutions was shown in [25, Theorem 3.1].

3.6 Existence for the entropy formulation

In this part we shortly recall the available existence theory for the Navier-Stokes-Fourier system in the entropy formulation (19 – 24). Unfortunately, to the best of our knowledge, there is only “analytical” proof without any numerical counterpart of the approximation scheme.

Constitutive relations

We start with a list of hypotheses imposed on the constitutive relations:

- In addition to Gibbs' equation (1), the pressure $p = p(\varrho, \vartheta)$ and the specific internal energy $e = e(\varrho, \vartheta)$ satisfy the *hypothesis of thermodynamic stability*

$$\frac{\partial p(\varrho, \vartheta)}{\partial \varrho} > 0, \quad c_v(\varrho, \vartheta) = \frac{\partial e(\varrho, \vartheta)}{\partial \vartheta} > 0 \quad (38)$$

for all $\varrho > 0$, $\vartheta > 0$, see [3].

- The internal energy and the pressure take the form

$$e(\varrho, \vartheta) = e_m(\varrho, \vartheta) + \frac{a}{\varrho} \vartheta^4, \quad p(\varrho, \vartheta) = p_m(\varrho, \vartheta) + \frac{a}{4} \vartheta^4, \quad a > 0, \quad (39)$$

where e_m , p_m represent molecular components augmented by radiation, see [28, Chapter 1]. Moreover, p_m and e_m satisfy the monoatomic gas equation of state

$$p_m(\varrho, \vartheta) = \frac{2}{3} e_m(\varrho, \vartheta). \quad (40)$$

In this context, Gibbs' equation (1) yields

$$p_m(\varrho, \vartheta) = \vartheta^{5/2} P\left(\frac{\varrho}{\vartheta^{3/2}}\right); \quad \text{whence } e_m(\varrho, \vartheta) = \frac{3}{2} \vartheta \frac{\vartheta^{3/2}}{\varrho} P\left(\frac{\varrho}{\vartheta^{3/2}}\right). \quad (41)$$

- Thermodynamics stability (38) implies

$$P(0) = 0, \quad P'(Z) > 0, \quad 0 < \frac{\frac{5}{3}P(Z) - P'(Z)Z}{Z} < c \text{ for any } Z > 0, \quad (42)$$

where, in addition, we require the specific heat at constant volume to be uniformly bounded.

- In accordance with (42), the function $Z \mapsto \frac{P(Z)}{Z}$ is non-increasing, and we suppose

$$\lim_{Z \rightarrow \infty} \frac{P(Z)}{Z^{5/3}} = p_\infty > 0. \quad (43)$$

- The transport coefficients $\mu = \mu(\vartheta)$, $\eta = \eta(\vartheta)$, and $\kappa = \kappa(\vartheta)$ in (27), (28) depend on the absolute temperature,

$$\underline{\mu}(1 + \vartheta^\alpha) \leq \mu(\vartheta) \leq \bar{\mu}(1 + \vartheta^\alpha), \quad |\mu'(\vartheta)| \leq c \text{ for all } \vartheta > 0, \quad \frac{2}{5} < \alpha \leq 1, \quad \underline{\mu} > 0, \quad (44)$$

$$0 \leq \mu(\vartheta) \leq \bar{\eta}(1 + \vartheta^\alpha) \text{ for all } \vartheta > 0, \quad (45)$$

and

$$\underline{\kappa}(1 + \vartheta^3) \leq \kappa(\vartheta) \leq \bar{\kappa}(1 + \vartheta^3) \text{ for all } \vartheta > 0, \quad \underline{\kappa} > 0. \quad (46)$$

Global existence

We report the following existence result proved in [28, Chapter 3, Theorem 3.1]:

Theorem 2. *Let $\Omega \subset \mathbb{R}^3$ be a bounded domain of class $C^{2+\nu}$. Suppose that the pressure p and the internal energy e are interrelated through (38–41), where $P \in C[0, \infty) \cap C^3(0, \infty)$ satisfies the structural hypotheses (42), (43). Let the transport coefficients μ, η, κ be continuously differentiable functions of the temperature ϑ satisfying (44–46). Let the initial data $\varrho_0, \vartheta_0, \mathbf{u}_0$ be given such that*

$$\varrho_0, \vartheta_0 \in L^\infty(\Omega), \varrho_0 > 0, \vartheta_0 > 0 \text{ a.a. in } \Omega, \mathbf{u}_0 \in L^2(\Omega; \mathbb{R}^3).$$

Then the Navier-Stokes-Fourier system (19–24) admits a weak solution $\varrho, \vartheta, \mathbf{u}$ in $(0, T) \times \Omega$ belonging to the class:

$$\varrho \geq 0 \text{ a.a. in } (0, T) \times \Omega,$$

$$\varrho \in C([0, T]; L^1(\Omega)) \cap L^\infty(0, T; L^{5/3}(\Omega)) \cap L^\beta((0, T) \times \Omega)$$

for a certain $\beta > \frac{5}{3}$;

$$\vartheta > 0 \text{ a.a. in } (0, T) \times \Omega, \vartheta \in L^\infty(0, T; L^4(\Omega)) \cap L^2(0, T; W^{1,2}(\Omega)),$$

$$\vartheta^3, \log(\vartheta) \in L^2(0, T; W^{1,2}(\Omega));$$

$$\mathbf{u} \in L^2(0, T; W_0^A(\Omega; \mathbb{R}^3)), \quad A = \frac{8}{5 - \alpha}, \quad \varrho \mathbf{u} \in C_{\text{weak}}(0, T; L^{5/4}(\Omega; \mathbb{R}^3)).$$

4 Relative energy, dissipative solutions, stability

In this section, we address the problem of *stability* of solutions to the Navier-Stokes-Fourier system. In particular, we find a convenient way how to measure the distance of a weak solution $[\varrho, \vartheta, \mathbf{u}]$ to an arbitrary trio of sufficiently smooth functions $[r, \Theta, \mathbf{U}]$. The hypothesis of thermodynamics stability (38) will play a crucial role.

4.1 Relative (modulated) energy

For an energetically isolated system, the total energy and mass

$$E = \int_{\Omega} \left[\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e(\varrho, \vartheta) \right] dx, \quad M = \int_{\Omega} \varrho dx,$$

are constants of motion, while the total entropy

$$S = \int_{\Omega} \varrho s(\varrho, \vartheta) \, dx$$

is non-decreasing in time. In particular, the ballistic free energy

$$H_{\Theta}(\varrho, \vartheta) = \varrho (e(\varrho, \vartheta) - \Theta s(\varrho, \vartheta)), \quad \Theta > 0 \text{ a positive constant,}$$

augmented by the kinetic energy gives rise to a *Lyapunov functional*

$$E_B = \int_{\Omega} \left[\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho (e(\varrho, \vartheta) - \Theta s(\varrho, \vartheta)) \right] \, dx.$$

As a direct consequence of the thermodynamic stability hypothesis (38), the ballistic free energy function enjoys two remarkable properties:

$$\varrho \mapsto H_{\Theta}(\varrho, \Theta) \text{ is convex;} \tag{47}$$

$$\vartheta \mapsto H_{\Theta}(\varrho, \vartheta) \text{ attains its global minimum at } \vartheta = \Theta.$$

This motivates the following definition of the *relative (modulated) energy functional*

$$\begin{aligned} \mathcal{E}(\varrho, \vartheta, \mathbf{u} | r, \Theta, \mathbf{U}) & \tag{48} \\ &= \int_{\Omega} \left[\frac{1}{2} \varrho |\mathbf{u} - \mathbf{U}|^2 + H_{\Theta}(\varrho, \vartheta) - \frac{\partial H_{\Theta}(r, \Theta)}{\partial \varrho} (\varrho - r) - H_{\Theta}(r, \Theta) \right] \, dx. \end{aligned}$$

It follows from (47) that

$$\mathcal{E}(\varrho, \vartheta, \mathbf{u} | r, \Theta, \mathbf{U}) \geq 0, \quad \mathcal{E}(\varrho, \vartheta, \mathbf{u} | r, \Theta, \mathbf{U}) = 0 \text{ only if } \varrho = r, \quad \vartheta = \Theta, \quad \mathbf{u} = \mathbf{U}.$$

4.2 Dissipative solutions

The strength of the concept of relative energy lies in the fact that the time evolution of $\mathcal{E}(\varrho, \vartheta, \mathbf{u} | r, \Theta, \mathbf{U})$ can be computed for any *weak* (entropy formulation) solution to the Navier-Stokes-Fourier system provided the trio of functions $[r, \Theta, \mathbf{U}]$ is smooth enough to be taken as admissible test functions in (19), (21), (23). Indeed the following *relative energy inequality*

$$\begin{aligned} & \left[\mathcal{E}(\varrho, \vartheta, \mathbf{u} | r, \Theta, \mathbf{U}) \right]_{t=0}^{t=\tau} & \tag{49} \\ & + \int_0^{\tau} \int_{\Omega} \frac{\Theta}{\vartheta} \left(\mathbb{S}(\vartheta, \nabla_x \mathbf{u}) : \nabla_x \mathbf{u} - \frac{\mathbf{q}(\vartheta, \nabla_x \vartheta) \cdot \nabla_x \vartheta}{\vartheta} \right) \, dx \, dt \\ & \leq \int_0^{\tau} \int_{\Omega} \varrho (\mathbf{u} - \mathbf{U}) \cdot \nabla_x \mathbf{U} \cdot (\mathbf{U} - \mathbf{u}) \, dx \, dt \\ & + \int_0^{\tau} \int_{\Omega} \varrho (s(\varrho, \vartheta) - s(r, \Theta)) (\mathbf{U} - \mathbf{u}) \cdot \nabla_x \Theta \, dx \, dt \end{aligned}$$

$$\begin{aligned}
 & + \int_0^\tau \int_\Omega \varrho \left(\partial_t \mathbf{U} + \mathbf{U} \cdot \nabla_x \mathbf{U} \right) \cdot (\mathbf{U} - \mathbf{u}) \, dx \, dt \\
 & + \int_0^\tau \int_\Omega (\mathbb{S}(\vartheta, \nabla_x \mathbf{u}) : \nabla_x \mathbf{U} - p(\varrho, \vartheta) \operatorname{div}_x \mathbf{U}) \, dx \, dt \\
 & - \int_0^\tau \int_\Omega \left(\varrho \left(s(\varrho, \vartheta) - s(r, \Theta) \right) \partial_t \Theta + \varrho \left(s(\varrho, \vartheta) - s(r, \Theta) \right) \mathbf{U} \cdot \nabla_x \Theta \right) \, dx \, dt \\
 & \quad - \int_0^\tau \int_\Omega \frac{\mathbf{q}(\vartheta, \nabla_x \vartheta)}{\vartheta} \cdot \nabla_x \Theta \, dx \, dt \\
 & + \int_0^\tau \int_\Omega \left(\left(1 - \frac{\varrho}{r} \right) \partial_t p(r, \Theta) - \frac{\varrho}{r} \mathbf{u} \cdot \nabla_x p(r, \Theta) \right) \, dx \, dt
 \end{aligned}$$

holds for any weak solution of the Navier-Stokes-Fourier system (19–24) and any trio of (smooth) test functions satisfying the compatibility conditions

$$r > 0, \Theta > 0, \mathbf{U}|_{\partial\Omega} = 0 \text{ or } \mathbf{U} \cdot \mathbf{n}|_{\partial\Omega} = 0 \text{ as the case may be,} \quad (50)$$

see [29, Section 3].

Following Lions [34], who proposed a similar definition for the incompressible Euler system, we may say that $[\varrho, \vartheta, \mathbf{u}]$ is a *dissipative solution* to the Navier-Stokes-Fourier system (3), (7), (11) if **(i)** $[\varrho, \vartheta, \mathbf{u}]$ belong to the regularity class specified in Theorem 2, **(ii)** $[\varrho, \vartheta, \mathbf{u}]$ satisfy the relative energy inequality (49) for any trio $[r, \Theta, \mathbf{U}]$ of sufficiently smooth (for all integrals in (49) to be well defined) test functions satisfying the compatibility conditions (50). As observed in [29, Section 3], any weak solution of the Navier-Stokes-Fourier system (19–24) is a dissipative solution. An existence theory in the framework of dissipative solutions was developed and applied to a vast class of physical spaces, including unbounded domains in R^3 , see Jesslé, Jin, Novotný [31].

4.3 Weak-strong uniqueness

An important property of the dissipative solutions is that they coincide with the strong solution of the same problem as long as the latter exists. Since weak solutions are dissipative, this remains true also for the weak solutions. The proof of the following statement can be found in [22, Theorem 6.2], and [29, Theorem 2.1]:

Theorem 3. *In addition to the hypotheses of Theorem 2, suppose that*

$$s(\varrho, \vartheta) = S \left(\frac{\varrho}{\vartheta^{3/2}} \right) + \frac{4a}{3} \frac{\vartheta^3}{\varrho}, \text{ with } S(Z) \rightarrow 0 \text{ as } Z \rightarrow \infty. \quad (51)$$

Let $\varrho, \vartheta, \mathbf{u}$ be a dissipative (weak) solution to the Navier-Stokes-Fourier system (19–24) in the set $(0, T) \times \Omega$. Suppose that the Navier-Stokes-Fourier system

admits a strong solution $\tilde{\varrho}, \tilde{\vartheta}, \tilde{\mathbf{u}}$ in the time interval $(0, T)$, emanating from the same initial data and belonging to the class

$$\partial_t \tilde{\varrho}, \partial_t \tilde{\vartheta}, \partial_t \tilde{\mathbf{u}}, \partial_x^m \tilde{\varrho}, \partial_x^m \tilde{\vartheta}, \partial_x^m \tilde{\mathbf{u}} \in L^\infty((0, T) \times \Omega), \quad m = 0, 1, 2.$$

Then

$$\varrho \equiv \tilde{\varrho}, \quad \vartheta \equiv \tilde{\vartheta}, \quad \mathbf{u} \equiv \tilde{\mathbf{u}}.$$

The extra hypothesis (51) reflects the Third law of thermodynamics and can be possibly relaxed. Whether or not the Navier-Stokes-Fourier system admits global-in-time strong solutions is an interesting open question, for small data results see Matsumura and Nishida [37], [38].

4.4 Weak solutions based on the thermal energy formulation

All results of this section so far applied to the entropy weak formulation (19–24) of the Navier-Stokes-Fourier system. A natural question to ask is to which extent the same idea may be used to the weak solution based on the thermal energy balance (25). As (23), (25) are apparently not equivalent in the weak framework, this is a non-trivial issue we want to address in this section. To this end, we consider smooth initial data, specifically,

$$\varrho_0, \vartheta_0 \in W^{3,2}(\Omega), \quad \mathbf{u}_0 \in W^{3,2}(\Omega; R^3), \quad \varrho_0, \vartheta_0 > 0. \quad (52)$$

Our first result provides necessary conditions for a weak solution of (19), (21), (25), and (26) to satisfy the entropy balance (23), see [20, Lemmas 2.3, 2.4].

Proposition 1. *Under the hypotheses of Theorem 1, let $[\varrho, \vartheta, \mathbf{u}]$ be a weak solution of the Navier-Stokes-Fourier system (19), (21), (25), (26) emanating from the initial data satisfying (52) and enjoying the extra regularity*

$$\varrho, \vartheta, \operatorname{div}_x \mathbf{u} \in L^\infty((0, T) \times \Omega), \quad \mathbf{u} \in L^\infty((0, T) \times \Omega; R^3).$$

Then

$$\varrho > 0, \quad \vartheta > 0 \quad \text{a.a. in } (0, T) \times \Omega$$

and the entropy balance (23) holds.

With the entropy inequality (23) at hand we may use the technique based on the relative energy functional $\mathcal{E}(\varrho, \vartheta, \mathbf{u} | r, \Theta, \mathbf{U})$ developed in the previous section. In particular, we have (see [20, Lemma 3.2]):

Proposition 2. *Under the hypotheses of Proposition 1, let $[\tilde{\varrho}, \tilde{\vartheta}, \tilde{\mathbf{u}}]$ be a strong solution of the Navier-Stokes-Fourier system defined in $(0, T) \times \Omega$, emanating from the initial data*

$$\tilde{\varrho}(0, \cdot) = \varrho(0, \cdot), \quad \tilde{\vartheta}(0, \cdot) = \vartheta(0, \cdot), \quad \tilde{\mathbf{u}}(0, \cdot) = \mathbf{u}(0, \cdot),$$

and belonging to the regularity class

$$\left\{ \begin{array}{l} \varrho, \vartheta \in C([0, T]; W^{3,2}(\Omega)), \mathbf{u} \in C([0, T]; W^{3,2}(\Omega; \mathbb{R}^3)), \\ \vartheta \in L^2(0, T; W^{4,2}(\Omega)), \mathbf{u} \in L^2(0, T; W^{4,2}(\Omega; \mathbb{R}^3)), \\ \partial_t \vartheta \in L^2(0, T; W^{2,2}(\Omega)), \partial_t \mathbf{u} \in L^2(0, T; W^{2,2}(\Omega; \mathbb{R}^3)). \end{array} \right\}$$

Then $[\tilde{\varrho}, \tilde{\vartheta}, \tilde{\mathbf{u}}]$ coincides with the weak solution $[\varrho, \vartheta, \mathbf{u}]$ in $(0, T) \times \Omega$.

Finally, we claim a conditional regularity result concerning the weak solutions of the Navier-Stokes-Fourier system, see [20, Theorem 2.2].

Theorem 4. *Under the hypotheses of Proposition 1, let $[\varrho, \vartheta, \mathbf{u}]$ be a weak solution of the Navier-Stokes-Fourier system (19), (21), (25), (26), emanating from regular initial data satisfying (52), and enjoying the extra regularity*

$$\varrho, \vartheta, \operatorname{div}_x \mathbf{u} \in L^\infty((0, T) \times \Omega), \mathbf{u} \in L^\infty((0, T) \times \Omega; \mathbb{R}^3).$$

Then $[\varrho, \vartheta, \mathbf{u}]$ is a strong (classical) solution of the problem in $(0, T) \times \Omega$.

4.5 Synergy analysis-numeric

From the practical point of view, the convergence of the numerical scheme established in Theorem 1 is not very satisfactory as it holds up to a suitable subsequence. As the weak solutions to the Navier-Stokes-Fourier system are not (known to be) unique, it is therefore not a priori excluded that there is another subsequence converging to a different solution of the same problem. However, combining Theorem 1 with Theorem 4 we may deduce the following *unconditional* convergence result that can be seen as an example of “synergy” between analysis and numerics:

Theorem 5. *Under the hypotheses of Theorem 1, let $[\varrho_h, \vartheta_h, \mathbf{u}_h]_{h>0}$ be a family of approximate solutions resulting from the numerical scheme (34 – 36), emanating from the initial data (52), such that*

$$\varrho_h > 0, \vartheta_h > 0,$$

and, in addition,

$$\varrho_h, \vartheta_h, |\mathbf{u}_h|, |\operatorname{div}_h \mathbf{u}_h| \leq M$$

a.a. in $(0, T) \times \Omega$ for a certain M independent of h .

Then

$$\varrho_h \rightarrow \varrho \text{ weakly-} (*) \text{ in } L^\infty(0, T; L^\gamma(\Omega)) \text{ and strongly in } L^1((0, T) \times \Omega),$$

$$\vartheta_h \rightarrow \vartheta \text{ weakly in } L^2(0, T; L^6(\Omega)),$$

$$\mathbf{u}_h \rightarrow \mathbf{u} \text{ weakly in } L^2(0, T; L^6(\Omega; \mathbb{R}^3)),$$

$$\nabla_h \mathbf{u}_h \rightarrow \nabla_x \mathbf{u} \text{ weakly in } L^2((0, T) \times \Omega; \mathbb{R}^{3 \times 3}),$$

where $[\varrho, \vartheta, \mathbf{u}]$ is the (strong) solution of the problem of the Navier-Stokes-Fourier system in $(0, T) \times \Omega$.

5 Viscosity solutions, inviscid limits

The Navier-Stokes-Fourier system describes the motion of a *viscous* and *heat* conducting fluid; the shear viscosity coefficient μ as well as the heat conductivity coefficient κ are (strictly) positive. Accordingly, the entropy production rate is strictly positive till the system reaches a thermodynamic equilibrium. The *inviscid* fluids, described by means of the Euler system, may be seen as the limit case of their viscous counterparts, where the viscosity and/or the heat conductivity vanishes. Solutions of the purely *hyperbolic* systems of conservation laws describing the motion of inviscid fluids exhibit very irregular behavior including the appearance of singularities - shock waves - in a finite time lap.

The concept of *weak* or even more general *measure-valued* solution is therefore indispensable in the mathematical theory of inviscid fluids. In the absence of a sufficiently strong dissipative mechanism, solutions of non-linear systems of conservation laws may develop fast oscillations and/or concentrations that inevitably give rise to singularities of various types. As shown in the nowadays classical work of Tartar [40], oscillations are involved in many problems, in particular in those arising in the context of inviscid fluids.

The well know deficiency of weak solutions is that they may not be uniquely determined in terms of the data and suitable admissibility criteria must be imposed in order to pick up the physically relevant ones, cf. Dafermos [12]. Although most of the admissibility constraints are derived from fundamental physical principles as the Second law of thermodynamics, their efficiency in eliminating the nonphysical solutions is still dubious, cf. Dafermos [13]. Recently, DeLellis and Székelyhidi [14] developed the method previously known as *convex integration* in the context of fluid mechanics, in particular for the Euler system. Among other interesting results, they show the existence of infinitely many solutions to the incompressible Euler system violating many of the standard admissibility criteria. Later, the method was adapted to the compressible case by Chiodaroli [9].

5.1 Euler-Fourier system

The class of weak solutions is apparently much larger than required by the classical theory. In other words, it might be easier to establish *existence* but definitely more delicate to show *uniqueness* among all possible weak solutions emanating from the same initial data. Adapting the technique of DeLellis and Székelyhidi [14] we show a rather illustrative but at the same time disturbing example of non-uniqueness in the context of fluid thermodynamics. To this end, consider the so-called Euler-Fourier system

$$\partial_t \varrho + \operatorname{div}_x(\varrho \mathbf{u}) = 0, \quad (53)$$

$$\partial_t(\varrho \mathbf{u}) + \operatorname{div}_x(\varrho \mathbf{u} \otimes \mathbf{u}) + \nabla_x(\varrho \vartheta) = 0, \quad (54)$$

$$\frac{3}{2} [\partial_t(\varrho\vartheta) + \operatorname{div}_x(\varrho\vartheta\mathbf{u})] - \Delta\vartheta = -\varrho\vartheta\operatorname{div}_x\mathbf{u}. \quad (55)$$

The system (53-55) can be viewed as a “special” case of the Navier-Stokes-Fourier system with $p = \varrho\vartheta$, $c_v = \frac{3}{2}$, $\mu = \eta = 0$, $\kappa = 1$. Although a correct physical justification of an inviscid, and, at the same time heat conducting fluid may be dubious, the system has been used as a suitable approximation in certain models, see Wilcox [41].

Infinitely many weak solutions

For the sake of simplicity, we consider the spatially periodic boundary conditions, meaning the underlying spatial domain

$$\Omega = \mathcal{T}^3 = ([-1, 1] |_{\{-1, 1\}})^3$$

is the “flat” torus.

We report the following result, see [11, Theorem 3.1]:

Theorem 6. *Let $T > 0$ be given. Let the initial data satisfy*

$$\varrho_0, \vartheta_0 \in C^3(\mathcal{T}^3), \quad \mathbf{u} \in C^3(\mathcal{T}^3; \mathbb{R}^3), \quad \varrho_0 > 0, \quad \vartheta_0 > 0 \text{ in } \mathcal{T}^3.$$

The initial-value problem for the Euler-Fourier system (53–55) admits infinitely many weak solutions in $(0, T) \times \Omega$ belonging to the class

$$\varrho \in C^2([0, T] \times \Omega), \quad \partial_t\vartheta \in L^p(0, T; L^p(\Omega)), \quad \nabla_x^2\vartheta \in L^p(0, T; L^p(\Omega; \mathbb{R}^{3 \times 3}))$$

for any $1 \leq p < \infty$,

$$\mathbf{u} \in C_{\text{weak}}([0, T]; L^2(\Omega; \mathbb{R}^3)) \cap L^\infty((0, T) \times \Omega; \mathbb{R}^3), \quad \operatorname{div}_x\mathbf{u} \in C^2([0, T] \times \Omega).$$

Infinitely many admissible weak solutions

The infinitely many solutions claimed in Theorem 6 are obtained in a non-constructive way by applying a variant of the method of convex integration in the spirit of DeLellis and Székelyhidi [14]. Apparently, many of them are non-physical since they violate the First law of thermodynamics, notably

$$\operatorname{ess\,lim\,inf}_{t \rightarrow 0^+} \int_{\Omega} \left[\frac{1}{2} \varrho |\mathbf{u}|^2 + c_v \varrho \vartheta \right] (t, \cdot) \, dx > \int_{\Omega} \left[\frac{1}{2} \varrho_0 |\mathbf{u}_0|^2 + c_v \varrho_0 \vartheta_0 \right] \, dx.$$

This fact motivates the following admissibility criterion:

We say that a weak solution $[\varrho, \vartheta, \mathbf{u}]$ of the Euler-Fourier system (53–55), supplemented with the initial data $[\varrho_0, \vartheta_0, \mathbf{u}_0]$, is admissible, if the energy inequality

$$\int_{\Omega} \left[\frac{1}{2} \varrho |\mathbf{u}|^2 + c_v \varrho \vartheta \right] (t, \cdot) \, dx \leq \int_{\Omega} \left[\frac{1}{2} \varrho_0 |\mathbf{u}_0|^2 + c_v \varrho_0 \vartheta_0 \right] \, dx$$

holds for a.a. $t \in [0, T]$.

Similarly to the Navier-Stokes-Fourier system, it can be shown that admissible solutions satisfy the relative energy inequality (49) (with $\mathbb{S} = 0$) and enjoy the weak-strong uniqueness property, cf. [11]. Still the following result holds true, see [11, Theorem 4.2]:

Theorem 7. *Let $T > 0$ and the initial data*

$$\varrho_0, \vartheta_0 \in C^3(\mathcal{T}^3), \quad \varrho_0 > 0, \quad \vartheta_0 > 0 \text{ in } \mathcal{T}^3$$

be given.

Then there exists

$$\mathbf{u}_0 \in L^\infty((0, T) \times \Omega; \mathbb{R}^3)$$

such that the initial-value problem for the Euler-Fourier system (53–55) admits infinitely many admissible weak solutions in $(0, T) \times \Omega$ belonging to the class specified in Theorem 6.

5.2 Riemann problem

At first glance, the initial velocity field $\mathbf{u}_0 \in L^\infty((0, T) \times \Omega; \mathbb{R}^3)$, the existence of which is claimed in Theorem 7, may seem rather irregular and possibly never reachable by trajectories emanating from “nice” initial data. Unfortunately, the situation is more delicate as illustrated by the following example due to Chiodaroli, DeLellis, and Kreml [10]. They consider the barotropic Euler system:

$$\partial_t \varrho + \operatorname{div}_x(\varrho \mathbf{u}) = 0, \tag{56}$$

$$\partial_t(\varrho \mathbf{u}) + \operatorname{div}_x(\varrho \mathbf{u} \otimes \mathbf{u}) + \nabla_x \varrho^2 = 0 \tag{57}$$

in $(0, T) \times \mathbb{R}^2$ endowed with the 1 – D Riemann initial data

$$\varrho_0 = \begin{cases} \varrho_- & \text{for } x_1 < 0, \\ \varrho_+ & \text{for } x_1 > 0 \end{cases} \tag{58}$$

$$u_0^1 = \begin{cases} v_- & \text{for } x_1 < 0, \\ v_+ & \text{for } x_1 > 0 \end{cases}, \quad u_0^2 = 0. \tag{59}$$

It can be shown (see Chiodaroli et al. [10, Theorem 1.1]) that there are initial data (58), (59) such that the Riemann problem (56–59) admits infinitely many weak solutions satisfying the standard entropy admissibility criterion

$$\partial_t \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho^2 \right) + \operatorname{div}_x \left[\left(\frac{1}{2} \varrho |\mathbf{u}|^2 + 2\varrho^2 \right) \mathbf{u} \right] \leq 0. \tag{60}$$

What is more, such solutions may be extended *backward in time* as Lipschitz functions yielding regular initial data for which system (56), (57), supplemented with (60), admits infinitely many solutions, see Chiodaroli et al. [10, Corollary 1.2]:

Theorem 8. *There exist Lipschitz initial data $[\varrho_0, \mathbf{u}_0]$ for which the barotropic Euler system (56), (57) admits infinitely many weak solutions in $R^2 \times \Omega$ satisfying (60). In addition, the initial data are independent of the x_2 variable and $u_0^2 = 0$.*

Riemann problem for the full Euler system

Motivated by the previous results we consider the Riemann problem for the full Euler system

$$\partial_t \varrho + \operatorname{div}_x(\varrho \mathbf{u}) = 0, \quad (61)$$

$$\partial_t(\varrho \mathbf{u}) + \operatorname{div}_x(\varrho \mathbf{u} \otimes \mathbf{u}) + \nabla_x(\varrho \vartheta) = 0, \quad (62)$$

$$\partial_t \left[\frac{1}{2} \varrho |\mathbf{u}|^2 + c_v \varrho \vartheta \right] + \operatorname{div}_x \left[\left(\frac{1}{2} \varrho |\mathbf{u}|^2 + c_v \varrho \vartheta + \varrho \vartheta \right) \mathbf{u} \right] = 0, \quad (63)$$

with the associated entropy inequality

$$\partial_t(\varrho s) + \operatorname{div}_x(\varrho s \mathbf{u}) \geq 0, \quad s = s(\varrho, \vartheta) \equiv \log \left(\frac{\vartheta^{c_v}}{\varrho} \right). \quad (64)$$

Similarly to the preceding section, we consider the Cauchy problem for the system (61 - 64) in the 2-D-case in the spatial domain

$$\Omega = R^1 \times \mathcal{T}^1, \quad \text{where } \mathcal{T}^1 \equiv [0, 1] \Big|_{\{0,1\}} \text{ is the "flat" sphere,}$$

meaning all functions of (t, x_1, x_2) are 1-periodic with respect to the second spatial coordinate x_2 .

We introduce 1-D Riemannian data

$$\varrho(0, x_1, x_2) = R_0(x_1), \quad R_0 = \begin{cases} R_L & \text{for } x_1 \leq 0, \\ R_R & \text{for } x_1 > 0, \end{cases} \quad (65)$$

$$\vartheta(0, x_1, x_2) = \Theta_0(x_1), \quad \Theta_0 = \begin{cases} \Theta_L & \text{for } x_1 \leq 0, \\ \Theta_R & \text{for } x_1 > 0, \end{cases} \quad (66)$$

$$u^1(0, x_1, x_2) = U_0(x_1), \quad U_0 = \begin{cases} U_L & \text{for } x_1 \leq 0, \\ U_R & \text{for } x_1 > 0, \end{cases}, \quad u^2(0, x_1, x_2) = 0. \quad (67)$$

As is well known, see for instance Chang and Hsiao [6], the Riemann problem (61 - 67) admits a solution

$$\varrho(t, x) = R(t, x_1) = R(\xi),$$

$$\vartheta(t, x) = \Theta(t, x) = \Theta(\xi),$$

$$\mathbf{u}(t, x) = [U(t, x), 0] = [U(\xi), 0]$$

depending solely on the self-similar variable $\xi = \frac{x_1}{t}$. Such a solution is unique in the class of BV solutions of the 1-D problem, see Chen and Frid [7], [8].

We focus on the class of Riemann data producing *shock-free solutions* (rarefaction waves), more specifically, solutions that are locally Lipschitz in the *open set* $(0, T) \times \Omega$. We claim the following, see [27, Theorem 2.1]:

Theorem 9. *Let $[\varrho, \vartheta, \mathbf{u}]$ be a weak solution of the Euler system (61 – 64) in $(0, T) \times \Omega$ originating from the Riemann data (65 - 67) and satisfying the associated far field conditions. Suppose in addition that the Riemann data (65 - 67) give rise to a shock-free solution $[R, \Theta, U]$ of the 1-D Riemann problem.*

Then

$$\varrho = R, \quad \vartheta = \Theta, \quad \mathbf{u} = [U, 0] \text{ a.a. in } (0, T) \times \Omega.$$

In the light of this result, we may conjecture that the possibility of infinitely many solutions provided by the method of convex integration occurs only if the weak solution dissipates mechanical energy. A definitive answer to this question, however, remains completely open.

5.3 Viscosity solutions

The aforementioned examples reopened the old problem of suitable admissibility criteria to be imposed on the weak solutions to inviscid fluid systems. A natural one, advocated by Bardos et al. [2], admits only those solutions obtained as an inviscid limit of the associated viscous flow represented in our framework by the Navier-Stokes-Fourier system. In certain cases, indeed, such a selection process may eliminate the “wild” solutions constructed by the method of convex integration.

To provide some support to this conjecture, we consider the barotropic Navier-Stokes system:

$$\partial_t \varrho + \operatorname{div}_x(\varrho \mathbf{u}) = 0, \quad (68)$$

$$\partial_t(\varrho \mathbf{u}) + \operatorname{div}_x(\varrho \mathbf{u} \otimes \mathbf{u}) + \nabla_x p(\varrho) = \operatorname{div}_x \mathbb{S}(\nabla_x \mathbf{u}), \quad (69)$$

supplemented with the constitutive relations for the pressure

$$p(\varrho) = a\varrho^\gamma, \quad a > 0, \quad \gamma > 1, \quad (70)$$

and the viscous stress

$$\mathbb{S}(\nabla_x \mathbf{u}) = \mu \left(\nabla_x \mathbf{u} + \nabla_x^t \mathbf{u} - \frac{2}{N} \operatorname{div}_x \mathbf{u} \mathbb{I} \right) + \eta \operatorname{div}_x \mathbf{u} \mathbb{I}, \quad \mu > 0, \quad \eta \geq 0, \quad (71)$$

along with its one-dimensional counterpart:

$$\partial_t R + \partial_y(RV) = 0, \quad (72)$$

$$\partial_t(RV) + \partial_y(RV^2) + \partial_y p(R) = \left[2\mu \left(1 - \frac{1}{N} \right) + \eta \right] \partial_{y,y}^2 V. \quad (73)$$

With the obvious identification $x_1 = y$, $\varrho(x) = R(x_1)$, $\mathbf{u}(x) = [V(x_1), 0, \dots, 0]$, any solution of problem (72), (73) satisfies also the extended system (68–71).

The 1D-dimensional fluid motion is nowadays well-understood, see Antontsev, Kazhikhov and Monakhov [1]. In particular, problem (72), (73) considered in the interval $(0, 1)$, and supplemented with the boundary conditions

$$V(t, 0) = V(t, 1) = 0, \quad t \in (0, T), \quad (74)$$

and the initial conditions

$$R(0, \cdot) = R_0 > 0, \quad V(0, \cdot) = V_0 \quad (75)$$

admits a (unique) weak solution for a fairly vast class of initial data, see Amosov and Zlotnik [42]. Moreover, the solutions are regular provided the initial data are smooth enough, see Kazhikhov [32].

Our goal is to show that, unlike their “Eulerian” counterparts discussed in the previous section, solutions of the 1D-problem (72), (73) are stable in the class of weak solutions system (68–71). To this end, we consider a domain $\Omega \subset R^N$, $N = 2, 3$,

$$\Omega = (0, 1) \times \mathcal{T}^{N-1}, \quad \text{where } \mathcal{T}^{N-1} \equiv ((0, 1)|_{\{0,1\}})^{N-1} \text{ is the torus in } R^{N-1},$$

specifically all functions defined in Ω are 1-periodic with respect to the variables x_j , $j > 1$. Accordingly, any solution r , V of problem (68–71) can be extended to be constant in x_j , $j > 1$.

We say that a pair of functions $[\varrho, \mathbf{u}]$ represent a *finite energy weak solution* to the Navier-Stokes system (68–71) in the space-time cylinder $(0, T) \times \Omega$, supplemented with the boundary conditions

$$\mathbf{u}|_{\partial\Omega} = 0, \quad (76)$$

and the initial conditions

$$\varrho(0, \cdot) = \varrho_0, \quad \mathbf{u}(0, \cdot) = \mathbf{u}_0 \quad (77)$$

if:

The density ϱ is a non-negative function, $\varrho \in C_{\text{weak}}([0, T]; L^\gamma(\Omega))$, $\mathbf{u} \in L^2(0, T; W_0^{1,2}(\Omega; R^N))$, $\varrho\mathbf{u} \in C_{\text{weak}}([0, T]; L^{2\gamma/(\gamma-1)}(\Omega))$, and

$$\int_{\tau_1}^{\tau_2} \int_{\Omega} [\varrho \partial_t \varphi + \varrho \mathbf{u} \cdot \nabla_x \varphi] = \left[\int_{\Omega} \varrho \varphi \, dx \right]_{t=\tau_1}^{t=\tau_2} \, dx dt \quad (78)$$

for any $0 \leq \tau_1 \leq \tau_2 \leq T$ and any $\varphi \in C^\infty([0, T] \times \overline{\Omega})$;

$$\int_{\tau_1}^{\tau_2} \int_{\Omega} [\varrho \mathbf{u} \cdot \partial_t \varphi + \varrho \mathbf{u} \otimes \mathbf{u} : \nabla_x \varphi + p(\varrho) \operatorname{div}_x \varphi - \mathbb{S}(\nabla_x \mathbf{u}) : \nabla_x \varphi] \, dx \, dt \quad (79)$$

$$= \left[\int_{\Omega} \varrho \mathbf{u} \cdot \varphi \, dx \right]_{t=\tau_1}^{t=\tau_2}$$

for any $0 \leq \tau_1 \leq \tau_2 \leq T$ and any $\varphi \in C_c^\infty([0, T] \times \Omega; \mathbb{R}^N)$.

In addition, the *energy inequality*

$$\int_{\Omega} \left[\frac{1}{2} \varrho |\mathbf{u}|^2 + P(\varrho) \right] (\tau, \cdot) \, dx \leq \int_{\Omega} \left[\frac{1}{2} \varrho_0 |\mathbf{u}_0|^2 + P(\varrho_0) \right] \, dx, \quad (80)$$

$$\text{with } P(\varrho) = \frac{a}{\gamma - 1} \varrho^\gamma,$$

holds for a.a. $\tau \in (0, T)$.

Finite energy weak solutions to the barotropic Navier-Stokes system are known to exist for any finite energy initial data whenever $\gamma > \frac{N}{2}$, see Lions [35] and [30].

We claim the following stability result, see [26, Theorem 2.1]:

Theorem 10. *Let*

$$\gamma > \frac{N}{2}, \quad q > \max \{2, \gamma'\}, \quad \frac{1}{\gamma} + \frac{1}{\gamma'} = 1 \text{ if } N = 2,$$

$$q > \max \left\{ 3, \frac{6\gamma}{5\gamma - 6} \right\} \text{ if } N = 3.$$

Let $[R, V]$ be a (strong) solution of the one-dimensional problem (72–75), with the initial data belonging to the class

$$R_0 \in W^{1,q}(0, 1), \quad R_0 > 0 \text{ in } [0, 1], \quad V_0 \in W_0^{1,q}(0, 1).$$

Let $[\varrho, \mathbf{u}]$ be a finite energy weak solution to the Navier-Stokes system (78–80) in $(0, T) \times \Omega$, supplemented with the conditions

$$\varrho_0 \in L^\infty(\Omega), \quad \varrho_0 > 0 \text{ a.a. in } \Omega; \quad \mathbf{u}_0 \in L^2(\Omega; \mathbb{R}^N).$$

Then

$$\int_{\Omega} \left[\frac{1}{2} \varrho |\mathbf{u} - \mathbf{V}|^2 + P(\varrho) - P'(R)(\varrho - R) - P(R) \right] (\tau, \cdot) \, dx \quad (81)$$

$$\leq c(T) \int_{\Omega} \left[\frac{1}{2} \varrho_0 |\mathbf{u}_0 - \mathbf{V}_0|^2 + P(\varrho_0) - P'(R_0)(\varrho_0 - R_0) - P(R_0) \right] \, dx$$

for a.a. $\tau \in (0, T)$.

We easily recognize a variant of the relative energy functional appearing in (81).

5.4 Vanishing dissipation limit for the Navier-Stokes-Fourier system

We conclude this part by a short discussion of the vanishing dissipation limit for the complete Navier-Stokes-Fourier system discussed in Section 3.6. To this end we suppose that the thermodynamics functions p , e , and s are given through (38–43), where the “radiation” coefficient $a > 0$ will be sent to zero in the asymptotic limit. More specifically, the target Euler system takes the form

$$\partial_t \varrho + \operatorname{div}_x(\varrho \mathbf{u}) = 0, \quad (82)$$

$$\partial_t(\varrho \mathbf{u}) + \operatorname{div}_x(\varrho \mathbf{u} \otimes \mathbf{u}) + \nabla_x p_m(\varrho, \vartheta) = 0, \quad (83)$$

$$\partial_t \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e_m(\varrho, \vartheta) \right) + \operatorname{div}_x \left[\left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e_m(\varrho, \vartheta) \right) \mathbf{u} + p_m(\varrho, \vartheta) \mathbf{u} \right] = 0, \quad (84)$$

considered on a bounded and smooth domain $\Omega \subset R^3$, and supplemented with the impermeability boundary condition

$$\mathbf{u} \cdot \mathbf{n}|_{\partial\Omega} = 0. \quad (85)$$

We remark that the total energy balance (84) can be equivalently reformulated as the entropy balance equation

$$\partial_t(\varrho s_m(\varrho, \vartheta)) + \operatorname{div}_x(\varrho s_m(\varrho, \vartheta) \mathbf{u}) = 0, \quad (86)$$

or the thermal energy balance

$$c_v(\varrho, \vartheta) (\partial_t(\varrho \vartheta) + \operatorname{div}_x(\varrho \vartheta \mathbf{u})) + \vartheta \frac{\partial p_m(\varrho, \vartheta)}{\partial \vartheta} \operatorname{div}_x \mathbf{u} = 0, \quad (87)$$

with

$$c_v(\varrho, \vartheta) = \frac{\partial e_m(\varrho, \vartheta)}{\partial \vartheta},$$

as long as the solution of the Euler system remains smooth.

A suitable existence result for the Euler system with the slip boundary condition (85) was obtained by Schochet [39, Theorem 1]. It asserts the local-in-time existence of a *classical* solution $[\varrho_E, \vartheta_E, \mathbf{u}_E]$ of the Euler system (82), (83), (85), (86) if:

- $\Omega \subset R^3$ is a bounded domain with a sufficiently smooth boundary, say $\partial\Omega$ of class C^∞ ;
- the initial data $[\varrho_{0,E}, \vartheta_{0,E}, \mathbf{u}_{0,E}]$ satisfy

$$\varrho_{0,E}, \vartheta_{0,E} \in W^{3,2}(\Omega), \quad \mathbf{u}_{0,E} \in W^{3,2}(\Omega; R^3), \quad \varrho_{0,E}, \vartheta_{0,E} > 0 \text{ in } \overline{\Omega}; \quad (88)$$

- the compatibility conditions

$$\partial_t^k \mathbf{u}_{0,E} \cdot \mathbf{n}|_{\partial\Omega} = 0 \quad (89)$$

hold for $k = 0, 1, 2$.

Navier-Stokes-Fourier system

We consider a slight modification of the Navier-Stokes-Fourier system, namely

$$\partial_t \varrho + \operatorname{div}_x(\varrho \mathbf{u}) = 0, \quad (90)$$

$$\partial_t(\varrho \mathbf{u}) + \operatorname{div}_x(\varrho \mathbf{u} \otimes \mathbf{u}) + \nabla_x p(\varrho, \vartheta) = \operatorname{div}_x \mathbb{S}(\vartheta, \nabla_x \mathbf{u}) - \lambda \mathbf{u} \quad (91)$$

$$\partial_t(\varrho s(\varrho, \vartheta)) + \operatorname{div}_x(\varrho s(\varrho, \vartheta) \mathbf{u}) + \operatorname{div}_x \left(\frac{\mathbf{q}}{\vartheta} \right) = \sigma, \quad \sigma = \frac{1}{\vartheta} \left(\mathbb{S}(\vartheta, \nabla_x \mathbf{u}) - \frac{\mathbf{q} \cdot \nabla_x \vartheta}{\vartheta} \right), \quad (92)$$

where $\mathbb{S}(\vartheta, \nabla_x \mathbf{u})$ is the viscous stress tensor given by Newton's law

$$\mathbb{S}(\vartheta, \nabla_x \mathbf{u}) = \nu \left[\mu(\vartheta) \left(\nabla_x \mathbf{u} + \nabla_x^t \mathbf{u} - \frac{2}{3} \operatorname{div}_x \mathbf{u} \mathbb{I} \right) + \eta(\vartheta) \operatorname{div}_x \mathbf{u} \mathbb{I} \right], \quad \nu > 0, \quad (93)$$

and $\mathbf{q} = \mathbf{q}(\vartheta, \nabla_x \vartheta)$ is the heat flux determined by Fourier's law

$$\mathbf{q} = -\omega \kappa(\vartheta) \nabla_x \vartheta, \quad \omega > 0. \quad (94)$$

The scaling parameters a , ν , ω , and λ are positive quantities supposed to vanish in the asymptotic limit. The momentum equation (91) contains an extra "damping" term $-\lambda \mathbf{u}$.

System (90–92) is supplemented by the complete slip boundary conditions

$$\mathbf{u} \cdot \mathbf{n}|_{\partial\Omega} = 0, \quad [\mathbb{S}(\vartheta, \nabla_x \mathbf{u}) \cdot \mathbf{n}] \times \mathbf{n}|_{\partial\Omega} = 0, \quad (95)$$

accompanied with the no-flux condition

$$\mathbf{q}(\vartheta, \nabla_x \vartheta) \cdot \mathbf{n}|_{\partial\Omega} = 0. \quad (96)$$

Relative energy inequality

Because of the presence of the extra term in the momentum equation (91), the relative energy inequality (49) takes the form

$$\begin{aligned} & \left[\mathcal{E} \left(\varrho, \vartheta, \mathbf{u} \middle| r, \Theta, \mathbf{U} \right) \right]_{t=0}^{t=\tau} \quad (97) \\ & + \int_0^\tau \int_\Omega \frac{\Theta}{\vartheta} \left(\mathbb{S}(\vartheta, \nabla_x \mathbf{u}) : \nabla_x \mathbf{u} - \frac{\mathbf{q}(\vartheta, \nabla_x \vartheta) \cdot \nabla_x \vartheta}{\vartheta} \right) dx dt + \lambda \int_0^\tau \int_\Omega |\mathbf{u}|^2 dx dt \\ & \leq \int_0^\tau \int_\Omega \varrho (\mathbf{u} - \mathbf{U}) \cdot \nabla_x \mathbf{U} \cdot (\mathbf{U} - \mathbf{u}) dx dt \\ & \quad + \int_0^\tau \int_\Omega \mathbb{S}(\vartheta, \nabla_x \mathbf{u}) : \nabla_x \mathbf{U} dx dt \\ & \quad - \int_0^\tau \int_\Omega \frac{\mathbf{q}(\vartheta, \nabla_x \vartheta)}{\vartheta} \cdot \nabla_x \Theta dx dt + \lambda \int_0^\tau \int_\Omega \mathbf{u} \cdot \mathbf{U} dx dt \end{aligned}$$

$$\begin{aligned}
 & + \int_0^\tau \int_\Omega \varrho \left(s(\varrho, \vartheta) - s(r, \Theta) \right) (\mathbf{U} - \mathbf{u}) \cdot \nabla_x \Theta \, dx \, dt \\
 & + \int_0^\tau \int_\Omega \varrho \left(\partial_t \mathbf{U} + \mathbf{U} \cdot \nabla_x \mathbf{U} \right) \cdot (\mathbf{U} - \mathbf{u}) \, dx \, dt - \int_0^\tau \int_\Omega p(\varrho, \vartheta) \operatorname{div}_x \mathbf{U} \, dx \, dt \\
 & - \int_0^\tau \int_\Omega \left(\varrho \left(s(\varrho, \vartheta) - s(r, \Theta) \right) \partial_t \Theta + \varrho \left(s(\varrho, \vartheta) - s(r, \Theta) \right) \mathbf{U} \cdot \nabla_x \Theta \right) \, dx \, dt \\
 & + \int_0^\tau \int_\Omega \left(\left(1 - \frac{\varrho}{r} \right) \partial_t p(r, \Theta) - \frac{\varrho}{r} \mathbf{u} \cdot \nabla_x p(r, \Theta) \right) \, dx \, dt
 \end{aligned}$$

for any trio of (smooth) test functions $[r, \Theta, \mathbf{U}]$ such that

$$r, \Theta > 0 \text{ in } \overline{\Omega}, \quad \mathbf{U} \cdot \mathbf{n}|_{\partial\Omega} = 0. \quad (98)$$

Similarly to the above, the relative energy inequality (97) for any weak solution $[\varrho, \vartheta, \mathbf{u}]$ of the Navier-Stokes-Fourier system specified through (19–23), where the total energy balance (24) is replaced by

$$\left[\int_\Omega \left[\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e(\varrho, \vartheta) \right] dx \right]_{t=\tau}^{t=0} + \lambda \int_0^\tau |\mathbf{u}|^2 \leq 0 \quad (99)$$

for a.a. $\tau \in [0, T]$.

Vanishing dissipation limit

The obvious idea how to compare a weak solution $[\varrho, \vartheta, \mathbf{u}]$ of the Navier-Stokes-Fourier system to the strong solution $\varrho_E, \vartheta_E, \mathbf{u}_E$ of the Euler system is to take the trio

$$r = \varrho_E, \quad \Theta = \vartheta_E, \quad \mathbf{U} = \mathbf{u}_E$$

as test functions in the relative energy inequality (97). Here we point out that such a step is essentially conditioned by our choice of the complete slip boundary condition (95) for the velocity field in the Navier-Stokes-Fourier system. Another choice of boundary behavior of \mathbf{u} , in particular the no-slip conditions (13), (14), would lead to the well known and sofar unsurmountable difficulties connected with the presence of a boundary layers, see the surveys of E [16] or Masmoudi [36].

We report the following result, see [23, Theorem 3.1]:

Theorem 11. *Let the following hypotheses be satisfied:*

- $\Omega \subset \mathbb{R}^3$ is a bounded domain with smooth boundary.
- The thermodynamic functions $p, e,$ and s are given by (39), (51), where p_m, e_m comply with (41–43), and, in addition,

$$P \in C^1[0, \infty) \cap C^5(0, \infty), \quad P'(Z) > 0 \text{ for all } Z \geq 0.$$

- The transport coefficients μ , η and λ are given by (44–46), with $\alpha = 1$.

Let $[\varrho_E, \vartheta_E, \mathbf{u}_E]$ be the classical solution of the Euler system (82–84), (85) in a time interval $(0, T)$, originating from the initial data $[\varrho_{0,E}, \vartheta_{0,E}, \mathbf{u}_{0,E}]$ satisfying (88), (89).

Finally, let $[\varrho, \vartheta, \mathbf{u}]$ be a weak solution of the Navier-Stokes-Fourier system (90–94), (95), (96), where the initial data $[\varrho_0, \vartheta_0, \mathbf{u}_0]$ satisfy

$$\varrho_0, \vartheta_0 > 0 \text{ a.a. in } \Omega,$$

$$\int_{\Omega} \varrho_0 \, dx \geq M, \quad \|\varrho_0\|_{L^\infty(\Omega)} + \|\vartheta_0\|_{L^\infty(\Omega)} + \|\mathbf{u}_0\|_{L^\infty(\Omega; \mathbb{R}^3)} \leq D,$$

and where the scaling parameters a , ν , ω , and λ are positive numbers.

Then

$$\begin{aligned} & \left[\mathcal{E} \left(\varrho, \vartheta, \mathbf{u} \mid \varrho_E, \vartheta_E, \mathbf{u}_E \right) \right]_{t=0}^{t=\tau} \\ & \leq c(T, M, D) \max \left\{ a, \nu, \omega, \lambda, \frac{\nu}{\sqrt{a}}, \frac{\omega}{a}, \left(\frac{a}{\sqrt{\nu^3 \lambda}} \right)^{1/3} \right\} \end{aligned}$$

for a.a. $\tau \in (0, T)$.

Corollary 1. Under the hypotheses of Theorem 11 suppose that

$$a, \nu, \omega, \lambda \rightarrow 0, \quad \text{and} \quad \frac{\omega}{a} \rightarrow 0, \quad \frac{\nu}{\sqrt{a}} \rightarrow 0, \quad \frac{a}{\sqrt{\nu^3 \lambda}} \rightarrow 0. \quad (100)$$

Then

$$\operatorname{ess\,sup}_{\tau \in (0, T)} \int_{\Omega} \left[\varrho |\mathbf{u} - \mathbf{u}_E|^2 + |\varrho - \varrho_E|^{5/3} + \varrho |\vartheta - \vartheta_E| \right] \, dx$$

$$\leq c(T, D, M) \times$$

$$\times \Lambda \left(a, \nu, \omega, \lambda, \|\varrho_0 - \varrho_{0,E}\|_{L^\infty(\Omega)}, \|\vartheta_0 - \vartheta_{0,E}\|_{L^\infty(\Omega)}, \|\mathbf{u}_0 - \mathbf{u}_{0,E}\|_{L^\infty(\Omega; \mathbb{R}^3)} \right),$$

where Λ is an explicitly computable function of its arguments,

$$\Lambda \left(a, \nu, \omega, \lambda, \|\varrho_0 - \varrho_{0,E}\|_{L^\infty(\Omega)}, \|\vartheta_0 - \vartheta_{0,E}\|_{L^\infty(\Omega)}, \|\mathbf{u}_0 - \mathbf{u}_{0,E}\|_{L^\infty(\Omega; \mathbb{R}^3)} \right) \rightarrow 0$$

provided a, ν, ω, λ satisfy (100), and

$$\|\varrho_0 - \varrho_{0,E}\|_{L^\infty(\Omega)}, \|\vartheta_0 - \vartheta_{0,E}\|_{L^\infty(\Omega)}, \|\mathbf{u}_0 - \mathbf{u}_{0,E}\|_{L^\infty(\Omega; \mathbb{R}^3)} \rightarrow 0.$$

The convergence result stated in Corollary 1 is *path dependent*, the parameters a, ν, ω, λ are interrelated through (100). It is easy to check that (100) holds provided, for instance,

$$a \rightarrow 0, \quad \nu = a^\alpha, \quad \omega = a^\beta, \quad \lambda = a^\gamma, \quad \text{where } \beta > 1, \quad \frac{1}{2} < \alpha < \frac{2}{3}, \quad 0 < \gamma < 1 - \frac{3}{2}\alpha.$$

6 Acknowledgement

The research of leading to these results has received funding from the European Research Council under the European Union's Seventh Framework Programme (FP7/2007-2013)/ ERC Grant Agreement 320078. The Institute of Mathematics of the Academy of Sciences of the Czech Republic is supported by RVO:67985840.

References

1. S. N. Antontsev, A. V. Kazhikhov, and V. N. Monakhov. *Krajevyye zadaci mekhaniki neodnorodnykh zidkostej*. Novosibirsk, 1983.
2. C. Bardos, M. C. Lopes Filho, Dongjuan Niu, H. J. Nussenzweig Lopes, and E. S. Titi. Stability of two-dimensional viscous incompressible flows under three-dimensional perturbations and inviscid symmetry breaking. *SIAM J. Math. Anal.*, **45**(3):1871–1885, 2013.
3. S. E. Bechtel, F.J. Rooney, and M.G. Forest. Connection between stability, convexity of internal energy, and the second law for compressible Newtonian fluids. *J. Appl. Mech.*, **72**:299–300, 2005.
4. F. Brezzi and M. Fortin. *Mixed and hybrid finite element methods*, volume 15 of *Springer Series in Computational Mathematics*. Springer-Verlag, New York, 1991.
5. H. Callen. *Thermodynamics and an Introduction to Thermostatistics*. Wiley, New Yoerk, 1985.
6. T. Chang and L. Hsiao. *The Riemann problem and interaction of waves in gas dynamics*, volume 41 of *Pitman Monographs and Surveys in Pure and Applied Mathematics*. Longman Scientific & Technical, Harlow; copublished in the United States with John Wiley & Sons, Inc., New York, 1989.
7. G.-Q. Chen and H. Frid. Uniqueness and asymptotic stability of Riemann solutions for the compressible Euler equations. *Trans. Amer. Math. Soc.*, **353**(3):1103–1117 (electronic), 2001.
8. G.-Q. Chen, H. Frid, and Y. Li. Uniqueness and stability of Riemann solutions with large oscillation in gas dynamics. *Comm. Math. Phys.*, **228**(2):201–217, 2002.
9. E. Chiodaroli. A counterexample to well-posedness of entropy solutions to the compressible Euler system. *J. Hyperbolic Differ. Equ.*, **11**(3):493–519, 2014.
10. E. Chiodaroli, C. DeLellis, and O. Kreml. Global ill-posedness of the isentropic system of gas dynamics. 2012. Preprint.
11. E. Chiodaroli, E. Feireisl, and O. Kreml. On the weak solutions to the equations of a compressible heat conducting gas. *Annal. Inst. Poincaré, Anal. Nonlinear.*, **32**:225–243, 2015.
12. C. M. Dafermos. *Hyperbolic conservation laws in continuum physics*. Springer-Verlag, Berlin, 2000.
13. C.M. Dafermos. The second law of thermodynamics and stability. *Arch. Rational Mech. Anal.*, **70**:167–179, 1979.
14. C. De Lellis and L. Székelyhidi, Jr. On admissibility criteria for weak solutions of the Euler equations. *Arch. Ration. Mech. Anal.*, **195**(1):225–260, 2010.

15. R.J. DiPerna and P.-L. Lions. Ordinary differential equations, transport theory and Sobolev spaces. *Invent. Math.*, **98**:511–547, 1989.
16. W. E. Boundary layer theory and the zero-viscosity limit of the Navier-Stokes equation. *Acta Math. Sin. (Engl. Ser.)*, **16**(2):207–218, 2000.
17. R. Eymard, T. Gallouët, and R. Herbin. Finite volume methods. In *Handbook of numerical analysis, Vol. VII*, Handb. Numer. Anal., VII, pages 713–1020. North-Holland, Amsterdam, 2000.
18. R. Eymard, T. Gallouët, R. Herbin, and J. C. Latché. A convergent finite element-finite volume scheme for the compressible Stokes problem. II. The isentropic case. *Math. Comp.*, **79**(270):649–675, 2010.
19. C. L. Fefferman. Existence and smoothness of the Navier-Stokes equation. In *The millennium prize problems*, pages 57–67. Clay Math. Inst., Cambridge, MA, 2006.
20. Feireisl and Y. Sun. Conditional regularity of very weak solutions to the Navier-Stokes-Fourier system. 2014. Preprint IM Prague.
21. E. Feireisl. *Dynamics of viscous compressible fluids*. Oxford University Press, Oxford, 2004.
22. E. Feireisl. Relative entropies in thermodynamics of complete fluid systems. *Discr. and Cont. Dyn. Syst. Ser. A*, **32**:3059–3080, 2012.
23. E. Feireisl. Vanishing dissipation limit for the Navier-Stokes-Fourier system. *Commun. Math. Sci.*, 2015. Submitted.
24. E. Feireisl, T. Karper, and Novotný. A convergent mixed numerical method for the Navier-Stokes-Fourier system. *IMA J. Numer. Anal.*, 2014. Submitted.
25. E. Feireisl, T. Karper, and A. Novotný. On a convergent numerical scheme for the full navier-stokes-fourier system. *IMA J. Numer. Math.*, 2014. Submitted.
26. E. Feireisl and O. Kreml. Uniqueness of rarefaction waves in multidimensional Euler system. *J. Hyperbolic Equations*, 2014. Submitted.
27. E. Feireisl, O. Kreml, and A. Vasseur. Stability of the isentropic Riemann solutions of the full multidimensional Euler system. *SIAM J. Math. Anal.*, 2015. To appear.
28. E. Feireisl and A. Novotný. *Singular limits in thermodynamics of viscous fluids*. Birkhäuser-Verlag, Basel, 2009.
29. E. Feireisl and A. Novotný. Weak-strong uniqueness property for the full Navier-Stokes-Fourier system. *Arch. Rational Mech. Anal.*, **204**:683–706, 2012.
30. E. Feireisl, A. Novotný, and H. Petzeltová. On the existence of globally defined weak solutions to the Navier-Stokes equations of compressible isentropic fluids. *J. Math. Fluid Mech.*, **3**:358–392, 2001.
31. D. Jesslé, B. J. Jin, and A. Novotný. Navier-Stokes-Fourier system on unbounded domains: weak solutions, relative entropies, weak-strong uniqueness. *SIAM J. Math. Anal.*, **45**(3):1907–1951, 2013.
32. A. V. Kazhikhov. Correctness “in the large” of mixed boundary value problems for a model system of equations of a viscous gas. *Dinamika Splošn. Sredy*, (Vyp. 21 Tecenie Zidkost. so Svobod. Granicami):18–47, 188, 1975.
33. J.-L. Lions. *Quelques méthodes de résolution des problèmes aux limites non linéaires*. Dunod, Gauthier - Villars, Paris, 1969.
34. P.-L. Lions. *Mathematical topics in fluid dynamics, Vol.1, Incompressible models*. Oxford Science Publication, Oxford, 1996.
35. P.-L. Lions. *Mathematical topics in fluid dynamics, Vol.2, Compressible models*. Oxford Science Publication, Oxford, 1998.

36. N. Masmoudi. Examples of singular limits in hydrodynamics. *In Handbook of Differential Equations, III, C. Dafermos, E. Feireisl Eds., Elsevier, Amsterdam, 2006.*
37. A. Matsumura and T. Nishida. The initial value problem for the equations of motion of viscous and heat-conductive gases. *J. Math. Kyoto Univ.*, **20**:67–104, 1980.
38. A. Matsumura and T. Nishida. The initial value problem for the equations of motion of compressible and heat conductive fluids. *Comm. Math. Phys.*, **89**:445–464, 1983.
39. S. Schochet. The compressible Euler equations in a bounded domain: Existence of solutions and the incompressible limit. *Commun. Math. Phys.*, **104**:49–75, 1986.
40. L. Tartar. Compensated compactness and applications to partial differential equations. *Nonlinear Anal. and Mech., Heriot-Watt Sympos., L.J. Knopps editor, Research Notes in Math 39, Pitman, Boston*, pages 136–211, 1975.
41. C. H. Wilcox. *Sound propagation in stratified fluids*. Appl. Math. Ser. 50, Springer-Verlag, Berlin, 1984.
42. A. A. Zlotnik and A. A. Amosov. Generalized solutions “in the large” of equations of the one-dimensional motion of a viscous barotropic gas. *Dokl. Akad. Nauk SSSR*, **299**(6):1303–1307, 1988.